Evolution of Langmuir waves in a plasma contaminated by variable-charge impurities

S. V. Vladimirov,* K. N. Ostrikov,† and M. Y. Yu

Institut fu¨r Theoretische Physik I, Ruhr-Universita¨t Bochum, D-44780 Bochum, Germany

L. Stenflo

Department of Plasma Physics, Umea˚ University, S-90187 Umea˚, Sweden

(Received 13 July 1998)

The propagation of Langmuir waves in nonisothermal plasmas contaminated by fine dust particles with variable charge is investigated for a self-consistent closed system. Dust charge relaxation, ionization, recombination, and collisional dissipation are taken into account. It is shown that the otherwise unstable coupling of the Langmuir and dust-charge relaxation modes becomes stable and the Langmuir waves are frequency downshifted. [S1063-651X(98)08612-7]

PACS number(s): 52.25.Vy, 52.35.Fp, 52.25.Kn

Charged impurities or dust grains are often found in space and laboratory plasmas $[1,2]$. The highly charged dust particles can significantly affect the system since they carry a considerable percentage of the total negative charge of the plasma. In fact, most plasma waves are to some degree affected by the dusts $[3-7]$. The variable dust charge also leads to a new plasma mode, usually referred to as the charge relaxation mode (CRM), involving dust-charge fluctuations [3,8] originating from the dust-charge variation determined by the instantaneous local electrostatic potential. In most earlier investigations on waves in dusty plasmas, the fact that dusty plasma models are almost always thermodynamically open was sidestepped by invoking unspecified sources or sinks, whose details are nevertheless important for a rigorous treatment of the problem.

In the absence of ionization, recombination, and collisional dissipation, it was found $[4]$ that Langmuir waves can be unstable because of a coupling to the CRM. In this Brief Report we reconsider the problem of linear Langmuir wave propagation in a nonisothermal dusty plasma with dustcharge variation as well as the collisional effects. The latter may be due to collisions between electrons and ions, neutral atoms, or other electrons, as well as the elastic and inelastic (the dust-charging) collisions between the electrons and the dust. These processes exist in most dusty plasmas, which are usually of low temperature and partially ionized. In fact, the dust-charge relaxation process is itself closely associated with ionization and recombination, which maintain the averaged background particle number densities self-consistently during the perturbations by acting as sources and sinks. They also define the equilibrium or steady state. Here we show

ostrikov@tp1.ruhr-uni-bochum.de

that these dissipative processes lead to a net damping of the Langmuir waves in typical dusty plasma systems.

We consider the propagation of linear Langmuir waves in a nonisothermal $(T_e \gg T_i)$, where T_e and T_i are the electron and ion temperatures) plasma. The size of the dust grains is assumed to be much less than the intergrain distance, the electron Debye radius, and the wavelength of the waves, so they can be treated as heavy point masses. The charge of a dust grain varies because of the microscopic electron and ion currents flowing into the grain according to the potential difference between the dust surface and the adjacent plasma. The dusts are treated as an immobile background since the time scale of charge variation is much smaller than that of the dust motion $\lfloor 8 \rfloor$.

The equations describing the propagation of Langmuir waves are

$$
\partial_t n_e + \mathbf{\nabla} \cdot (n_e \mathbf{v}_e) = S,\tag{1}
$$

$$
\partial_t \mathbf{v}_e + \nu_{\text{eff}} \mathbf{v}_e = (e/m_e) \nabla \varphi - \gamma V_{Te}^2 \nabla n_e / n_{0e} ,\qquad (2)
$$

$$
\nabla^2 \varphi = -4\pi e (Z_i n_i - n_e - Z_d n_d), \qquad (3)
$$

where φ is the electrostatic potential and m_i , n_j (including the stationary value n_{0j}), and \mathbf{v}_i are the mass, density, and fluid velocity of the species $j = e$, *i*, and *d* for electron, ion, and dust, respectively. Furthermore, $Z_i e$ and $-Z_d e$ are the charges of the ions and dusts, γ is the adiabatic constant of the electrons, and $V_{Te} = (T_e / m_e)^{1/2}$ is the electron thermal velocity. We have also defined $S = -v_{ed}n_e + v_i n_e - \beta n_e^2$ $+\beta_{si}n_e^2 + \nabla \cdot (D_a \nabla n_e)$, where ν_{ed} is the collection rate of plasma electrons by the dust grains, v_i is the ionization rate, β is the volume recombination rate, β_{si} is the stepwise ionization rate, and D_a is the ambipolar diffusion coefficient. We have also defined the effective electron collision frequency $v_{\text{eff}} = v_e + v_e^{\text{el}} + v_e^{\text{ch}}$, where v_e is the rate of electron collisions with the neutral atoms and plasma particles, v_e^{el} is the rate of elastic (Coulomb) electron-dust collisions, and v_e^{ch} is the effective rate of collection of plasma particles by the dust.

We shall assume that the wave perturbations behave like \sim exp[$i(kz - \omega t)$] and use the probe model for dust charging.

^{*}Present address: Research Center for Theoretical Astrophysics, School of Physics, The University of Sydney, New South Wales 2006, Australia. Electronic address:

s.vladimirov@physics.usyd.edu.au

[†] Permanent address: Kharkov State University and Scientific Center for Physical Technologies, 2 Novgorodskaya #93, 310145 Kharkov, Ukraine. Electronic address:

In the absence of the perturbations, the system is quasineutral so that $Z_i n_{i0} = n_{e0} + Z_d n_{d0}$. The dust-charge relaxation process is described by the charge balance equation $[8]$

$$
d_t q_d = I_e(q_d) + I_i(q_d),\tag{4}
$$

where q_d is the average charge on the dust grain and $I_e(q_d)$ and $I_i(q_d)$ are the electron and ion grain currents flowing into the grain surface. The quantities q_d , I_e , and I_i involve both steady-state and perturbed components, i.e., $q_d = q_{d0}$ $+q_{d1}$, $I_e = I_{e0} + I_{e1}$, and $I_i = I_{i0} + I_{i1}$, where

$$
I_{e0} = -\pi a^2 e (8T_e / \pi m_e)^{1/2} n_{e0} \exp[e \Delta \varphi_g / T_e],
$$
 (5)

$$
I_{i0} = \pi a^2 e Z_i (8T_i / \pi m_i)^{1/2} n_{i0} [1 - e \Delta \varphi_g / T_i]
$$
 (6)

are the steady-state electron and ion currents at the grain surface and $q_{d0} = C\Delta\varphi$ _g is the stationary charge of the grain. Here $a \left(\ll r_{De} \right)$ is the grain radius, r_{De} is the electron Debye radius, $C = a(1 + a/r_{\text{De}})$ is the effective grain capacitance, and $\Delta \varphi_g = \varphi_g - \varphi_0$ is the steady-state potential difference between the grain and the adjacent plasma. The floating potential φ_0 is determined by equating the equilibrium electron and ion currents (5) and (6) .

For Langmuir waves, we can neglect I_{i1} with respect to I_{e1} since it is on the slower ion time scale. From Eqs. (4)–(6) we obtain for the perturbed dust charge

$$
d_t q_{d1} + \nu_{ch} q_{d1} = -|I_{e0}| n_{e1} / n_{e0}, \qquad (7)
$$

where $v_{\text{ch}} = a \omega_{pi}^2 A / \sqrt{2 \pi} V_{Ti}$ is the charging rate of the dust particle [8], defined by the equilibrium electron and ion microscopic currents (5) and (6) . Here V_{Ti} is the ion thermal velocity, $\tau = T_i / T_e$, $\mathcal{A} = 1 + \tau + \mathcal{Z}$, and $\mathcal{Z} = Z_d e^2 / a T_e$. The effective charging rate is $v_e^{\text{ch}} = v_{\text{ch}}P(4+\mathcal{Z})(\tau+\mathcal{Z})/A\mathcal{Z}$ and the electron capture rate at the grain surface is $[8,9]$ v_{ed} $= v_{ch}P(\tau + \mathcal{Z})/\mathcal{AZ}, P = Z_{d}n_{d0}/n_{e0}$. The frequency of elastic electron-dust collisions is $[9]$ $v_e^{\text{el}} = 4\sqrt{2\pi Z_d^2 n_d e^4 \Lambda}$ /3 $m_e^2 V_{Te}^3$, where $\Lambda = \ln(r_{De}/a)$ is the Coulomb logarithm. The expressions for the rate v_e of electron collisions as well as for v_i , β , β_{si} , and D_a can be found in Ref. [10]. Equations (1)–(3) and (7) describe the coupling between the high-frequency electrostatic Langmuir waves and the CRM.

To determine the stationary electron plasma density we assume that the pressure is not too low, such that recombination losses prevail over diffusion losses. The last term in *S* can then be ignored. We then obtain from Eq. (1) the lowestorder (steady-state) electron plasma density $n_{e0} = (v_i)$ $-\nu_{ed}/\beta_{\text{eff}}$, where $\beta_{\text{eff}} = \beta - \beta_{si}$. We note that the ionization rate must be high enough such that $v_i > v_{ed}$; otherwise no self-consistent stationary state exists.

Linearizing with respect to the wave perturbations, we obtain from Eq. (1)

$$
\partial_t n_{e1} + n_{e0} \nabla \cdot \mathbf{v}_e = - \nu_{ed} |_{Z_{d0}} n_{e1},
$$
\n
$$
- n_{e0} (\partial_{Z_d} \nu_{ed})_{Z_{d0}} Z_{d1} - 2 \beta_{eff} n_{e0} n_{e1} + \xi_1 \nu_i n_{e1},
$$
\n(8)

where n_{e1} and Z_{d1} are the perturbations of the electron density and dust charge, respectively, and ξ_1 depends on the model for direct ionization. We shall consider two models, namely, that of a density-fluctuation independent ionization rate, with $\xi_1 = 0$, and that of density-fluctuation-dependent ionization rate, with $\xi_1 = 1$.

From Eq. (3) one can easily obtain the perturbed electron density $n_{e1} = -(1/4\pi e)k^2\varphi + Z_{d1}n_{d0}$. Thus we have for the grain charge and electron density variations

$$
q_{d1} = ik^2 |I_{e0}| \phi / 4 \pi e n_{e0} (\omega + i \nu_{ch}^*)
$$
 (9)

and

$$
n_{e1} = \nabla (\tilde{\epsilon}_d \nabla \varphi) / 4 \pi e, \qquad (10)
$$

where $v_{ch}^* = v_{ch} + \tilde{\nu}$, $\tilde{\nu} = n_{d0} |I_{e0}| / n_{e0} e$, and $\tilde{\epsilon}_d = 1 - i \tilde{\nu} / (\omega)$ $+i\nu_{\rm ch}^{*}$). Furthermore, from Eq. (8) we have

$$
\frac{n_{e1}}{n_{e0}} = \frac{i}{\eta} \bigg[-(\partial_{Z_d} \nu_{ed})_{Z_{d0}} Z_{d1} + \frac{ie k^2 \varphi}{m_e(\omega + i \nu_{\text{eff}})} \bigg],\qquad(11)
$$

where $\eta = \omega - i \nu_{ed} + i \xi_2 \nu_i - \gamma k^2 V_{Te}^2 / (\omega + i \nu_{eff})$ and $\xi_2 = 2$ $-\xi_1$. Finally, equating Eqs. (11) [after substituting of Z_{d1} from Eq. (9) and (10) , we obtain the dispersion relation of the Langmuir waves

$$
D(\omega, k) = i \tilde{\nu} \frac{\omega + i \nu_{\text{eff}}}{\omega + i \nu_{\text{ch}}^*} \left[\eta + i \frac{n_{e0}}{n_{d0}} (\partial_{Z_d} \nu_{ed})_{Z_{d0}} \right], \qquad (12)
$$

where $D(\omega, k) = (\omega + i\nu_{\text{eff}})\eta - \omega_{pe}^2$ and ω_{pe} is the electron plasma frequency. This is the equation describing the linear coupling of the high-frequency Langmuir plasma waves with the CRM mode $\omega = -i v_{ch}^{\prime}$. It can be solved numerically for any given set of parameters. We also note that if v_e^{el} , v_e^{ch} , v_{ed} , and $(\partial_{Z_d} v_{ed})_{Z_{d0}}$ are set to zero in Eq. (12), one recovers the coupling equation of Ref. $[4]$.

It is instructive to estimate the effect of electron capture (by the dust grain) and dissipative collisions on the Langmuir waves. For this purpose it is convenient to make the ansatz $\omega \ge \nu_{\text{eff}}$, ν_{ed} , ν_{ch}^* . Setting $\omega = \omega_1 + \delta_1' + i \delta_1''$, where $\omega_1^2 = \omega_{pe}^2 + \gamma k^2 V_{Te}^2$, we find from Eq. (12)

$$
\delta_1' = -\left(\omega_{pe}^2 \tilde{\nu} \omega_1^2 - \mathcal{B}\right) \tilde{\nu} / 2\omega_1,\tag{13}
$$

where $B = a^3 n_{i0} \omega_{pi}^2 A/v_{ch}$ and

$$
2\,\delta_1'' = \omega_{pe}^2\,\tilde{\nu}/\omega_1^2 - \nu_{\rm eff} - \xi_2\,\nu_i + \nu_{ed} + \tilde{\nu}\,\nu_{\rm eff}\mathcal{B}/\omega_1^2,\qquad(14)
$$

so that the frequency of the Langmuir waves is down shifted and the waves are damped by most of the collisional processes included here. In the absence of the latter, the first term in Eq. (14) remains and it leads to the Langmuir wave instability discussed earlier [4]. Although the rate v_{ed} of electron capture by the dust also has a positive sign, it is always smaller than the term $-\xi_2 \nu_i$ involving ionization because $v_i > v_{ed}$ (required by the existence condition for the stationary state) and ξ_2 > 1.

For the CRM, we set $\omega = -i v_{ch}^* + i \delta_2$. From Eq. (12) one then obtains

$$
\delta_2 \omega^2 / \tilde{\nu} = 3k^2 V_{Te}^2 + (\nu_{ch}^* - \nu_{eff}) (\nu_{ch}^* + \xi_2 \nu_i - \nu_{ed} - \mathcal{B}),
$$
\n(15)

which shows that the dust charging rate is slightly reduced by the coupling with Langmuir waves.

We now estimate the average dust charge and the dissipation parameters v_{eff} , v_{ed} , v_{ch}^* , and $\tilde{\nu}$ for typical dusty plasmas. The factor $e\Delta\varphi_g/T_e$, which defines the average charge on a dust grain and can strongly affect the density ratio n_{d0}/n_{e0} through the quasineutrality condition, can be found from the condition of zero total current flowing into the dust in the absence of the high-frequency perturbations. For a typical dusty argon plasma, we have $T_e \sim 10$ eV, T_i \sim 1 eV, $r \sim$ 5 μ m, $n_{e0} \sim$ 5×10¹⁰ cm⁻³, and $n_{i0}/n_{e0} = 10$. One then obtains $e\Delta\varphi_g/T_e=-1.71$, $Z_{d0}=-6.12\times10^4$, and $n_{d0}/n_{e0} \approx 1.74 \times 10^{-4}$. One can also show that the inequality $\vec{v} \leq \omega$, where $\vec{v} = v_{\text{eff}}$, v_{ed} , v_{ch}^* , \vec{v} represents the dissipative effects invoked here, is satisfied. In fact, we find $\bar{\nu}$ (3×10⁷)–(5×10⁸) sec⁻¹ and ω ~10¹⁰ sec⁻¹, which validate our ansatz.

It is also necessary to verify the existence condition v_i $> v_{ed}$ for the stationary state with the equilibrium density $n_{e0} = (v_i - v_{ed})/\beta_{\text{eff}}$. To estimate the ionization frequency v_i we use the expression (10.9) of Ref. $[11]$ for the ionization rate (averaged over a Maxwellian distribution) $\langle \sigma_{\text{ion}} v_e \rangle$ in hydrogen, where σ_{ion} is the ionization crosssection. Accordingly, we have

$$
\nu_{i} = \frac{2 \times 10^{-7} N_{n}}{6.0 + T_{e}/U_{i}} \sqrt{\frac{T_{e}}{U_{i}}} \exp\left(-\frac{U_{i}}{T_{e}}\right),
$$
 (16)

where U_i is the ionization energy and N_n is the number density (in cm⁻³) of the neutral particles. We see that v_i is sensitive to the electron temperature. The threshold temperature T_e^{thres} may be estimated by setting $v_i = v_{ed}$. For the typical hydrogen plasma parameters $n_{i0} \sim 10^{10} \text{ cm}^{-3}$, $N_n \sim 10^{14}$ cm⁻³, $r \sim 1$ μ m, $P=1$, and $Z \sim 2$, we find $T_e^{\text{thres}} \sim 7$ eV. A

decrease of T_e would lead to a decrease of the ionization frequency and hence a violation of the existence condition for the stationary state. We note that T_e^{thres} depends weakly on the ion density. Furthermore, the threshold is affected by the degree of ionization (an increase of n_{i0} / N_n would lead to a downshift of T_e^{thres}), the dust size and charge (T_e^{thres} increases with *a* and Z_{d0}), and the number of electrons in the system (T_e^{thres} increases when n_{e0}/n_{i0} decreases). It should be pointed out that in using the expression (10.9) of Ref. [11], one should check the relation between T_e and U_{eff} $= (2/3)(U_i - U_*)$, where U_* is the energy of the first excited level. For electron temperatures exceeding U_{eff} , multistep ionization dominates and the term U_i in the exponent of Eq. (16) must be replaced by the lower valued U_* . In this case the direct-ionization approximation is no longer valid and the existence threshold for the stationary state may be lowered. For argon plasmas the corresponding expressions for the ionization frequencies are more complicated, but the exponential dependence on $-U_{(*,i)}/T_e$ remains the same [10]. Thus for argon plasmas similar results can be expected.

In conclusion, we have shown that if ionization, recombination, and other collisions are included, the linear coupling of the CRM and Langmuir waves leads to a damping and a frequency downshift of the waves. This result differs considerably from that where a uniform source is invoked to replace the electrons and ions lost to dust charging [4]. Thus the actual ionization and recombination processes that maintain the total charge balance of a dusty plasma system may be important in investigations of instabilities in dusty plasmas.

K.N.O. and S.V.V. are grateful to the Alexander von Humboldt Foundation for financial support. This work was supported by the Sonderforschungsbereich 191 Niedertemperatur Plasmen.

- $[1]$ C. K. Goertz, Rev. Geophys. **27**, 271 (1989).
- [2] A. Bouchoule and L. Boufendi, Plasma Sources Sci. Technol. **3**, 262 (1994).
- [3] S. V. Vladimirov, Phys. Rev. E **50**, 1422 (1994).
- [4] J. X. Ma and M. Y. Yu, Phys. Rev. E 50, R2431 (1994).
- [5] F. Verheest, Space Sci. Rev. **77**, 267 (1996), and the reference therein.
- [6] V. N. Tsytovich, Usp. Fiz. Nauk **40** 53 (1997) [Phys. Usp. **40**, 53 (1997)].
- [7] K. N. Ostrikov, M. Y. Yu, and N. A. Azarenkov, Phys. Rev. E

58, 2431 (1998).

- [8] V. N. Tsytovich and O. Havnes, Comments Plasma Phys. Control. Fusion **15**, 267 (1993).
- @9# S. V. Vladimirov and V. N. Tsytovich, Phys. Rev. E **58**, 2415 $(1998).$
- [10] L. M. Biberman, V. S. Vorob'ev, and I. T. Yakubov, *Kinetics of Non-Equilibrium Low-Temperature Plasmas* (Nauka, Moscow, 1982).
- [11] R. J. Goldston and P. H. Rutherford, *Introduction to Plasma Physics* (IOP Publishing, London, 1995).